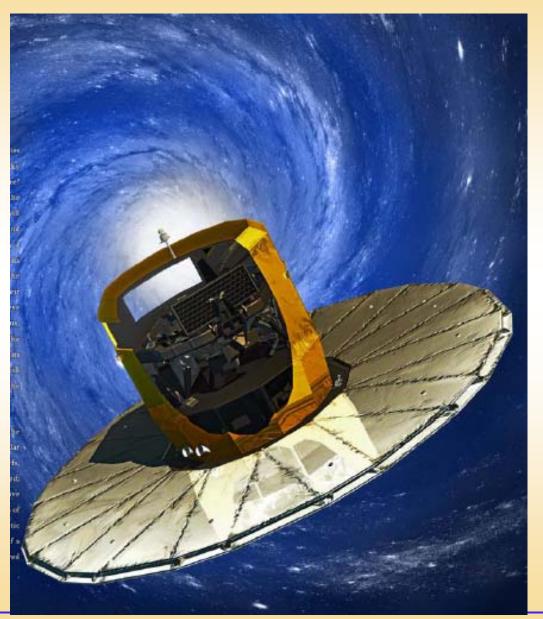


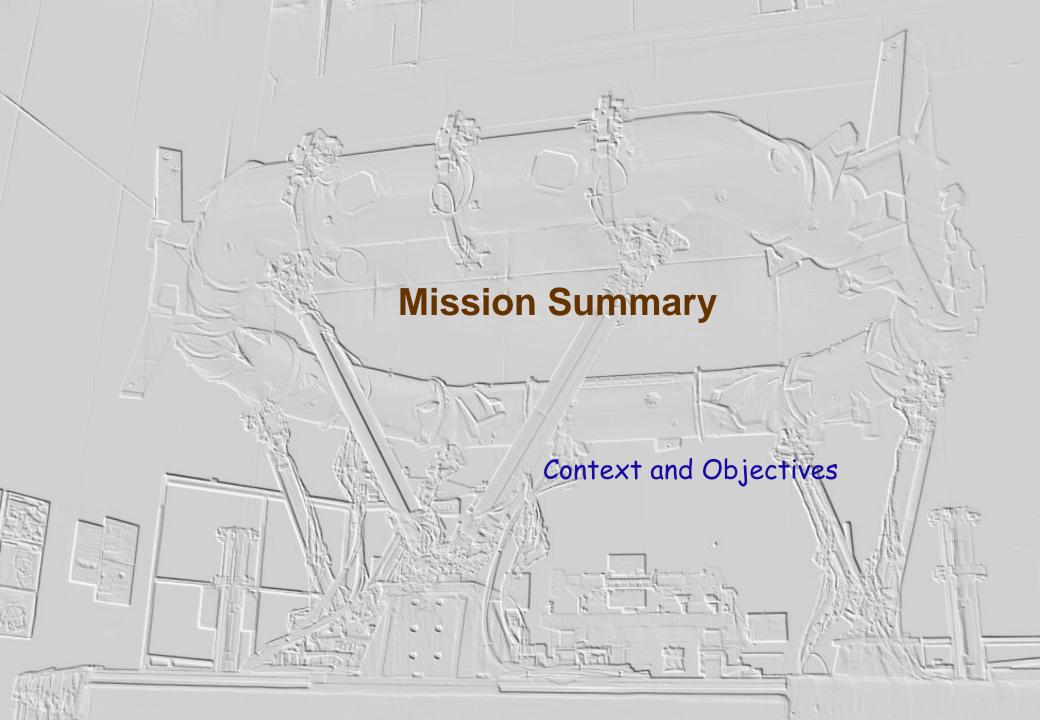


Outline



- Mission context
- Principles
- Gaia Performances
- Data releases





GAIA

109 stars

25 μ as @ V = 15 mag

ESA mission

Launch: 2013

Mission: 5 yrs

Photometry (~ 25 bands)

Radial velocity

Low resolution spectroscopy



Goals of Space Astrometry



- Primary Objectives not achievable from Earth
 - Ascertain the distances of the stars
 - absolute stellar parallaxes for astronomers
 - * Define and materialise the inertial frame
 - now based on extragalactic sources

■ Secondary objectives

- Astrophysics with astrometry, photometry, spectroscopy
 - stellar and galactic physics
 - detection of extrasolar planets
 - solar system dynamics
- Tests of fundamental physics in space
 - based on light path geometry



Space Astrometry: Past & Present



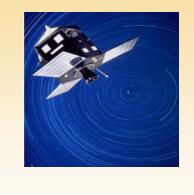
- A successful forerunner: HIPPARCOS (ESA)
 - accuracy of 1 mas ~ a coin @ 1000 km
- The unfortunate followers
 - accuracy of 0.1 mas ~ a nail @ 1000 km
 - * Roemer, FAME-1, FAME-2, DIVA, Lomonossov, AMEX

ESA US US DE RU US



- JASMINE (Japan) in the IR
- Cancelled (Dec 2010)
 - SIM (US) with 1 µas accuracy
- Funded launch 2011 2014
 - NanoJasmine [4 mas], J-MAPS (US) [1mas]
 - * Gaia (ESA): 25 μas (a hairwidth @ 1000 km)











Mission requirements summary



- A Stereoscopic Census of Our Galaxy
- Astrometry (V < 20):
 - completeness to 20 mag (on-board detection) 109 stars
 - parallax accuracy: 7 μas at <10 mag; 12-25 μas at 15 mag 100-300 μas at 20 mag
- Photometry (V < 20):
 - astrophysical diagnostics (low-dispersion photometry) + chromaticity
 - ◆ 8-20 mmag at 15 mag: Teff ~ 200 K, log g, [Fe/H] to 0.2 dex, extinction
- Radial velocity (V < 16.5-17):
 - Third component of space motion, perspective acceleration
 - <1 km/s at 13-13.5 mag and <15 km/s at 16.5-17 mag</p>





Assets of Gaia



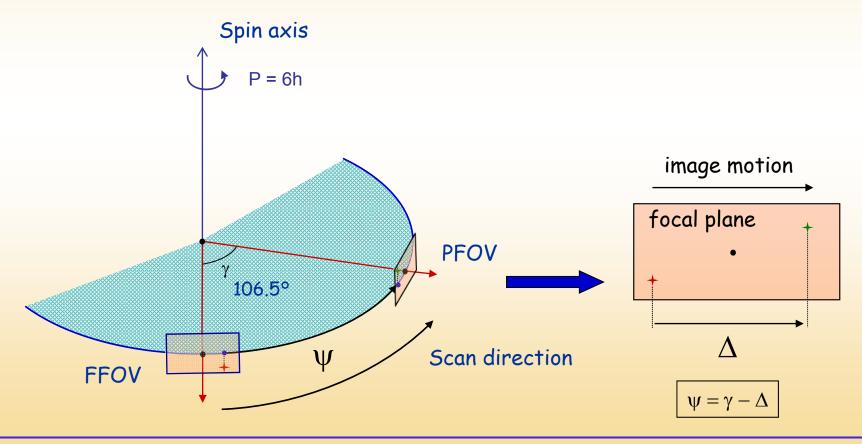
- A single mission with three nearly synchronous data taking
 - Astrometric, photometric and spectroscopic data
- GAIA is a scanning mission
 - no pointing, no change in the schedule Uniform coverage of the sky
- Quasi regular time sampling over 5 years
 - → 80 observations → photometry, orbits of binaries, asteroids
- Survey mission sensitivity limited
- Internal and autonomous detection system to G = 20
- Global astrometry of staggering precision
 - Internal metrology, thermal and mechanical stability
- Experienced and motivated community in Europe after Hipparcos
 - scientific and in industry



Global astrometry in space



- Wide angle measurements
- Two fields of view
- One common focal plane

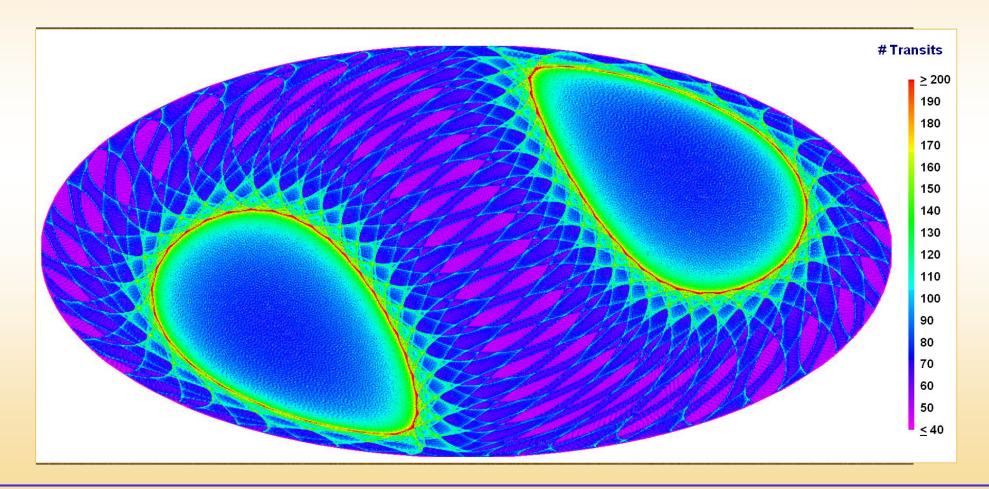




Scanning & Sky coverage



- Time average is a combination of the sky distribution and the scanning law
 - two different symmetries: galactic plane and eclitpic plane

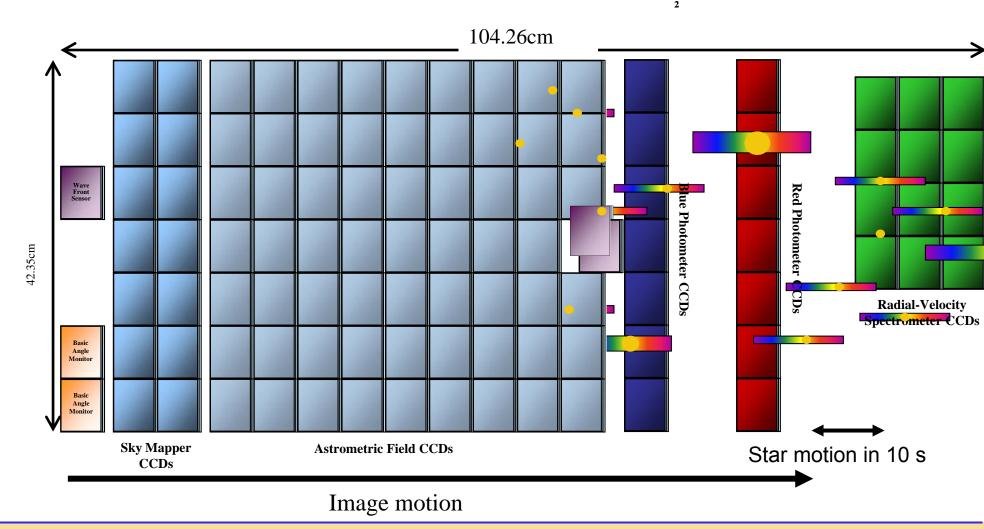




Multiplexing observations



106 CCDs, 938 million pixels, 2800 cm²

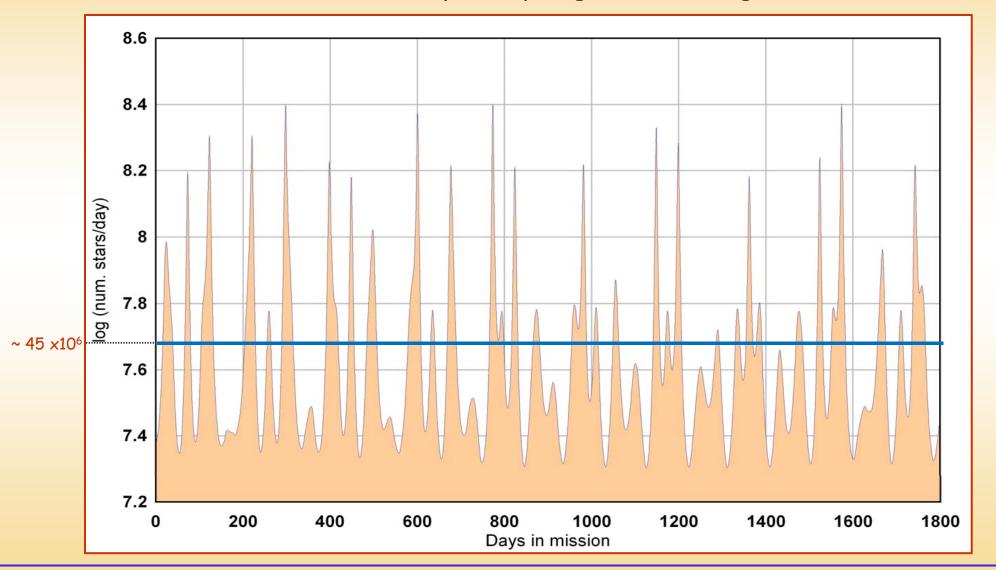




Number of sources per day



■ Number of sources detected per day (log scale) during the mission

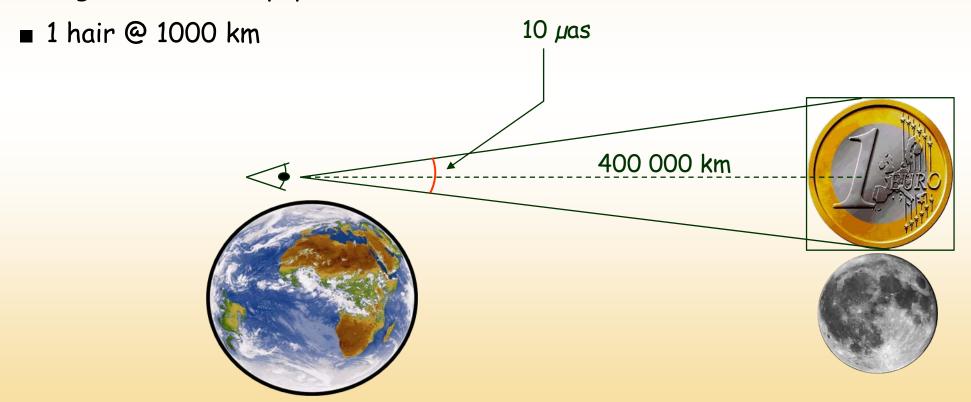




Performances 10 µas → Incredibly small!



- 0.3 mm displacement on the Earth
- Displacement of a 100 mas/yr star in one hour
- Motion of a fast moving minor planet in 100 μ s
- edge-on sheet of paper @ 2000 km





Astrometric accuracy: single transit



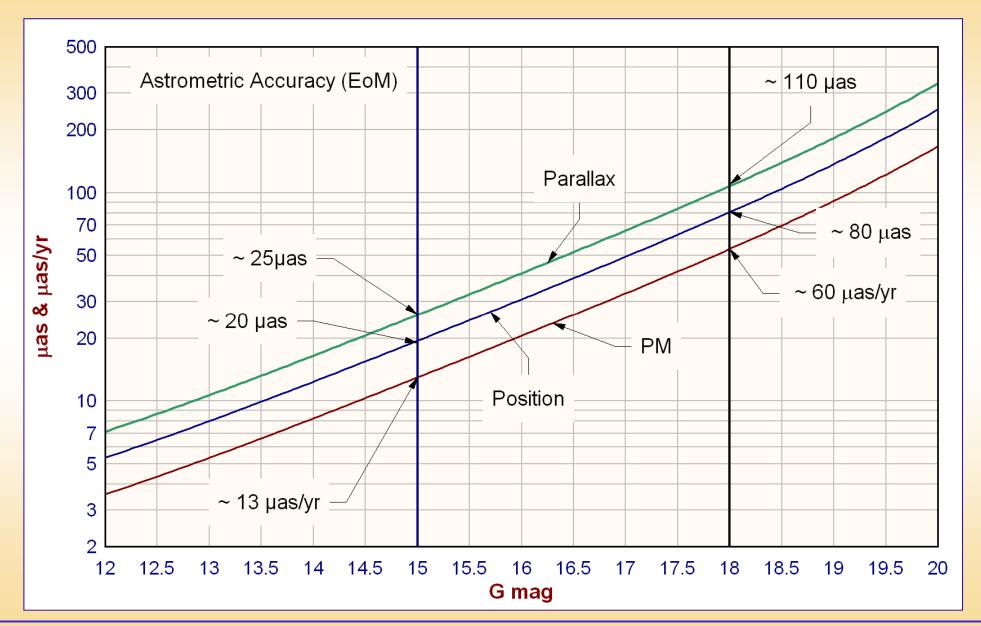
- Single observation accuracy → orbit, solar system
 - one field transit: integration over 9 AF CCDs
 - point source, 1D astrometry





Astrometric Accuracy: EOM



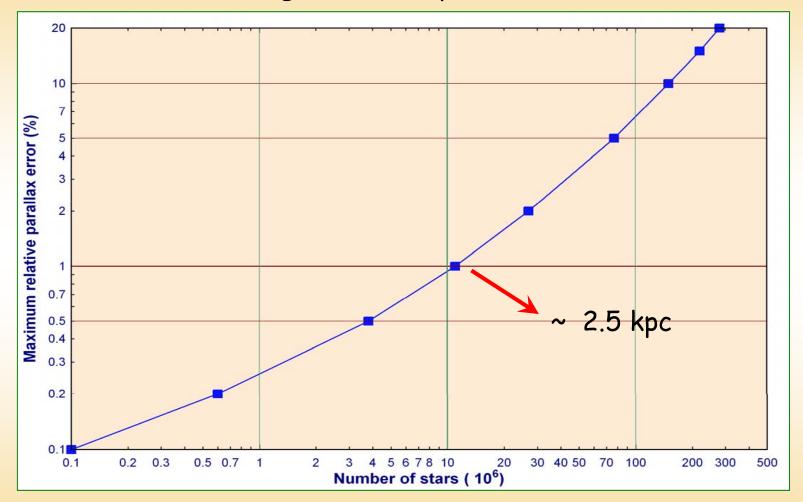




Distances for stellar physics



Accurate distances through the Galaxy



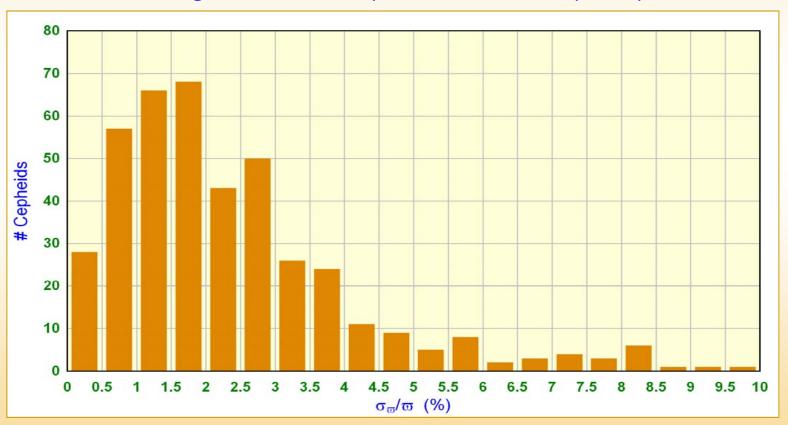
Recall: Hipparcos : 20,000 stars with σ_{π}/π < 10%



Cepheids with Gaia



- 15 d < 0.5 kpc, 65 d < 1 kpc, 165 d < 2 kpc
 - bright enough (V < 14)
- In the plot: 400 galactic cepheids from David Dunlap DB
 - \bullet distance and magnitude \rightarrow Gaia predicted accuracy for parallax

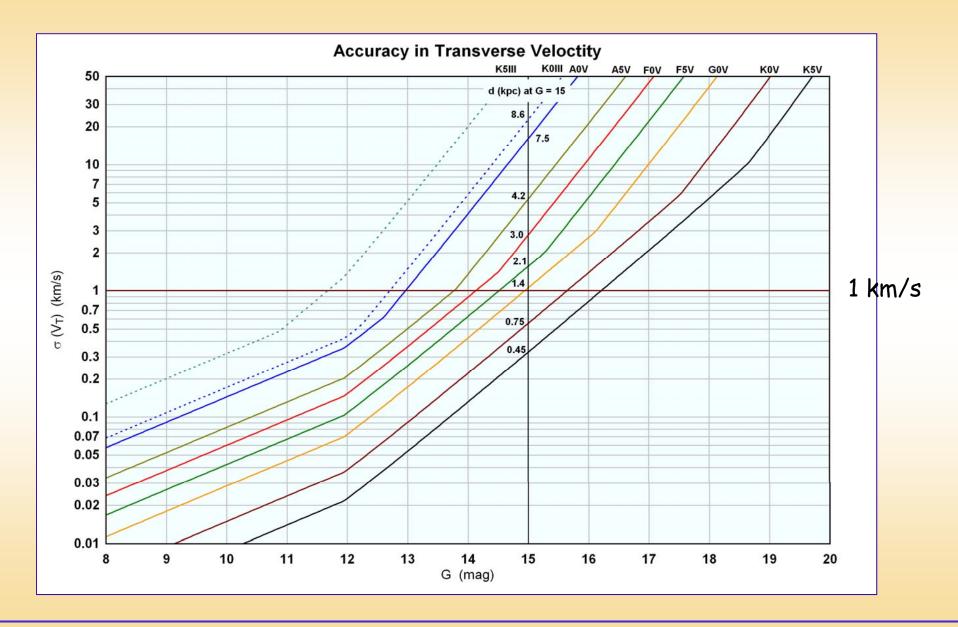


F. Mignard 2002, 2009



Transverse velocity estimate with Gaia

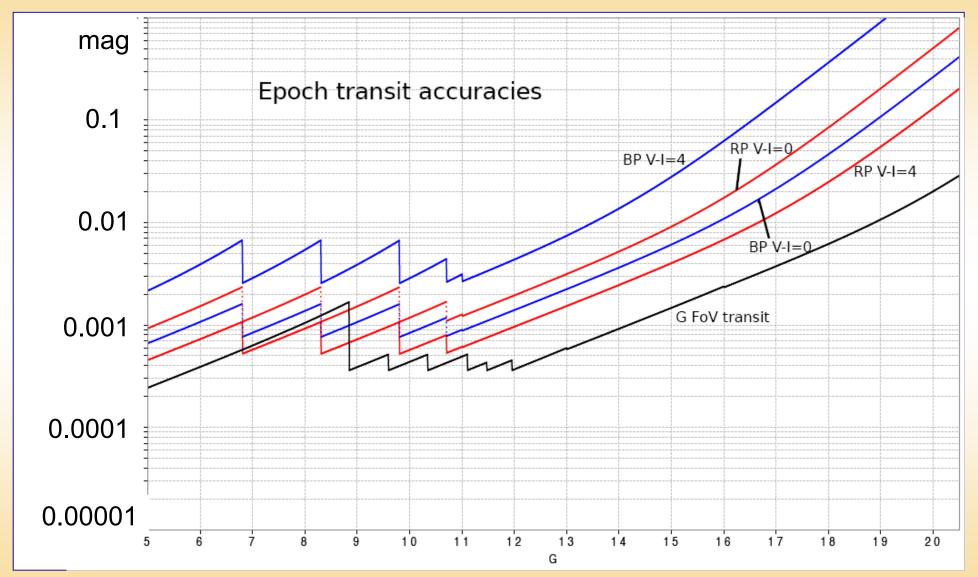






Photometric Performance over a transit





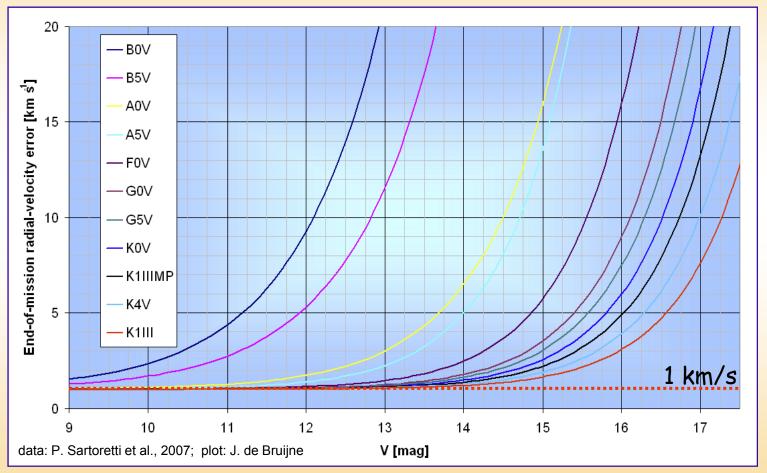
Credit: D. Evans



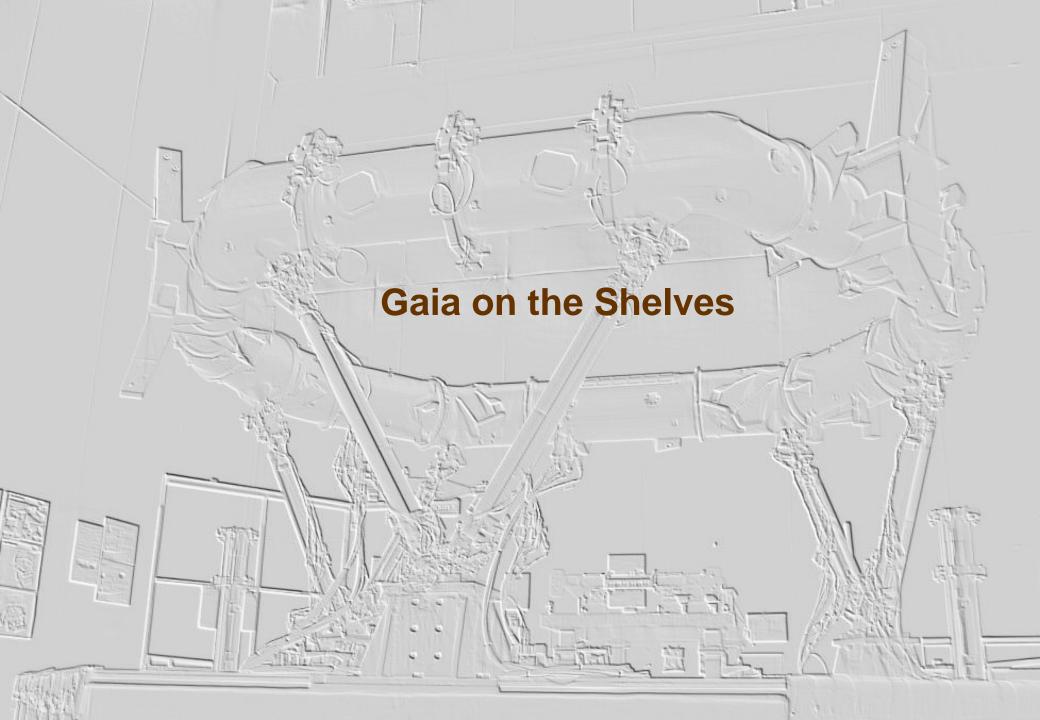
Radial velocity accuracy (EOM, km/s)



- Performances strongly dependent on stellar type
- Average of 40 transits (i.e 120 CCD crossings)



RAVE : $<V_r> ~ 2 \text{ km/s}, 9 < I < 12$





'Hardware ' already manufactured

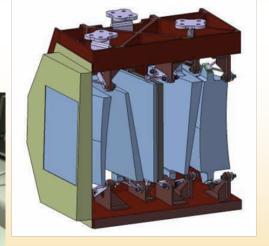




口













Gaia: Brased torus in place December 2009

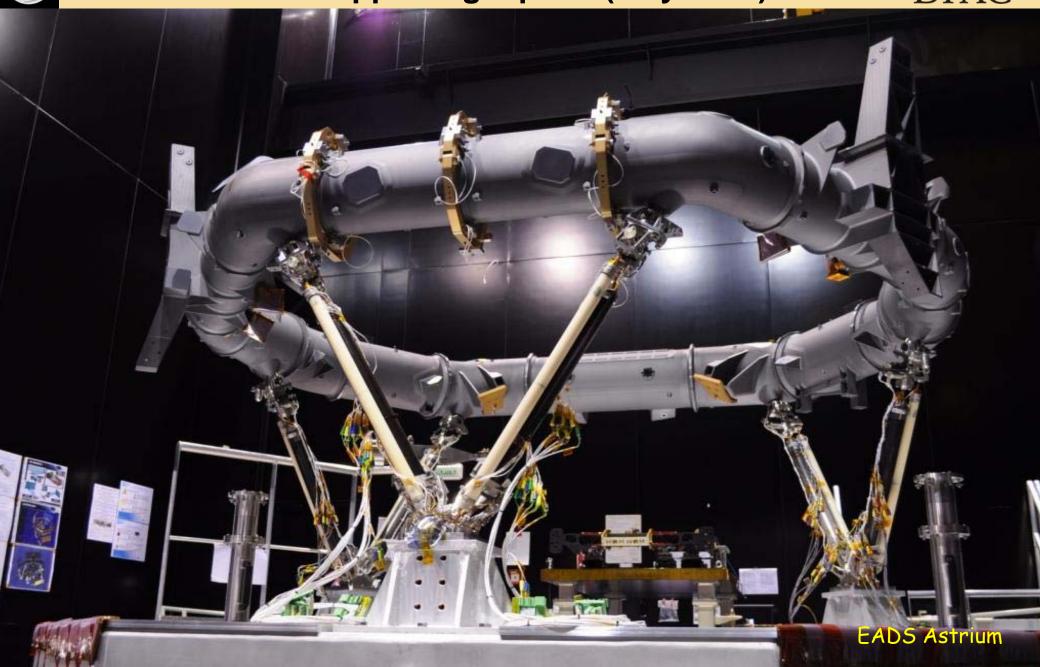






Torus with supporting bipods (July 2010)



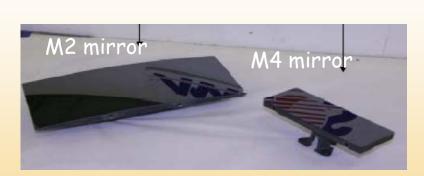




Gaia in construction early 2011







Credit: EADS Astrium









■ Nine out of the ten mirrors have been delivered



Credit: EADS Astrium, Boostec





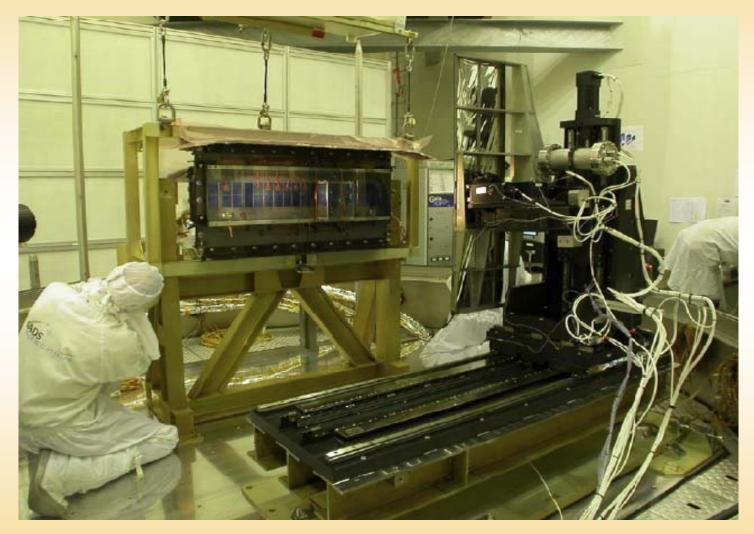
■ Rehearsal of the M1 mounting (March 2011)





Testing CCD acquisition mode



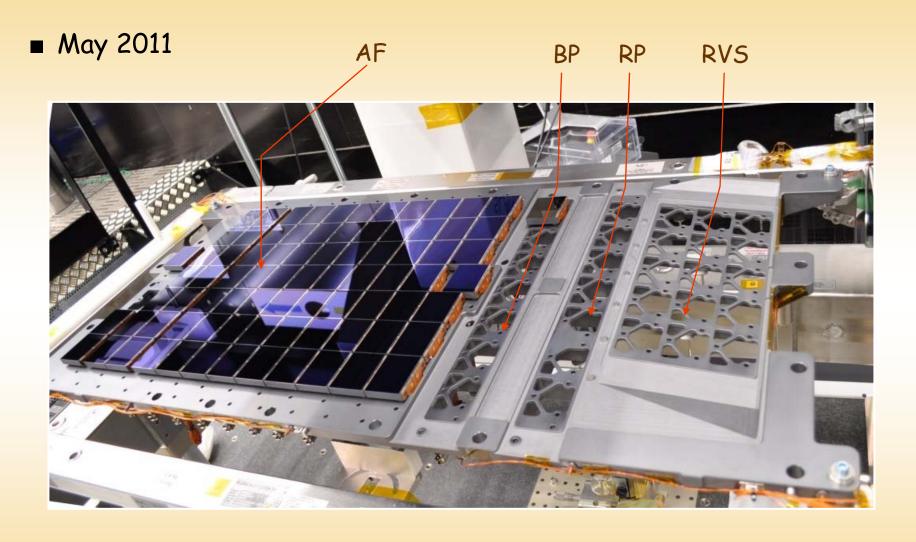


EADS Astrium



CCD integration







June: CCD integration completed!



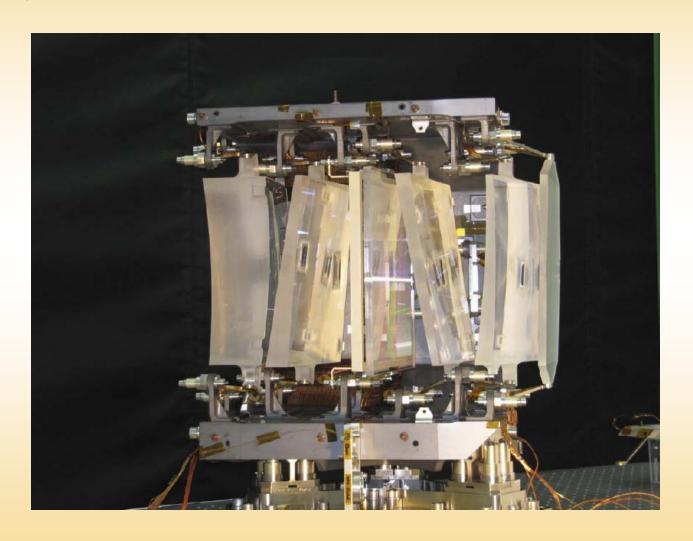




RVS Integrated



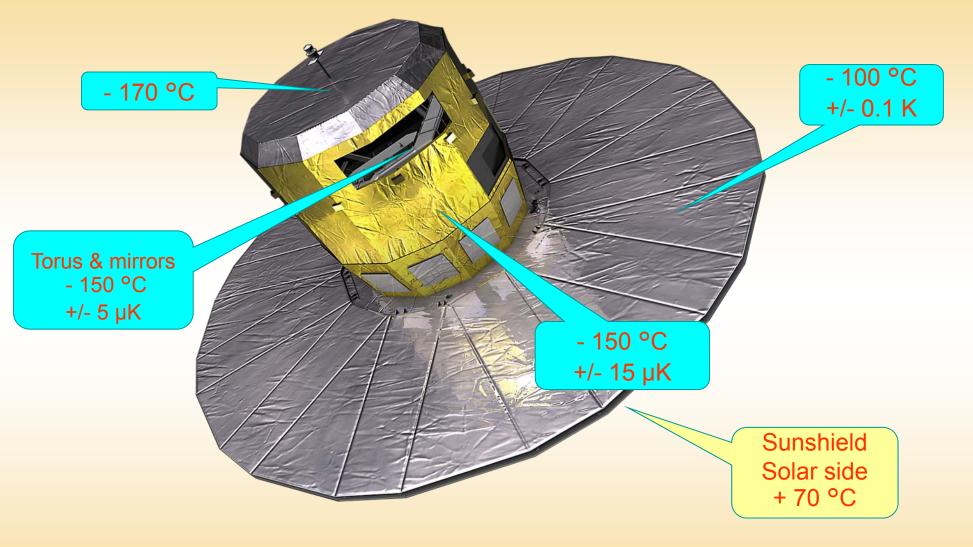
■ May - June 2011





Thermal Insulation







Credit: F. Chassat, Astrium



The Sun-Shield



First full deployement test with new motor





Soyuz Launchpad near Kouru







Preparations for the simulated first Soyuz flight



■ Pictures taken on the 29/04 and 02/05

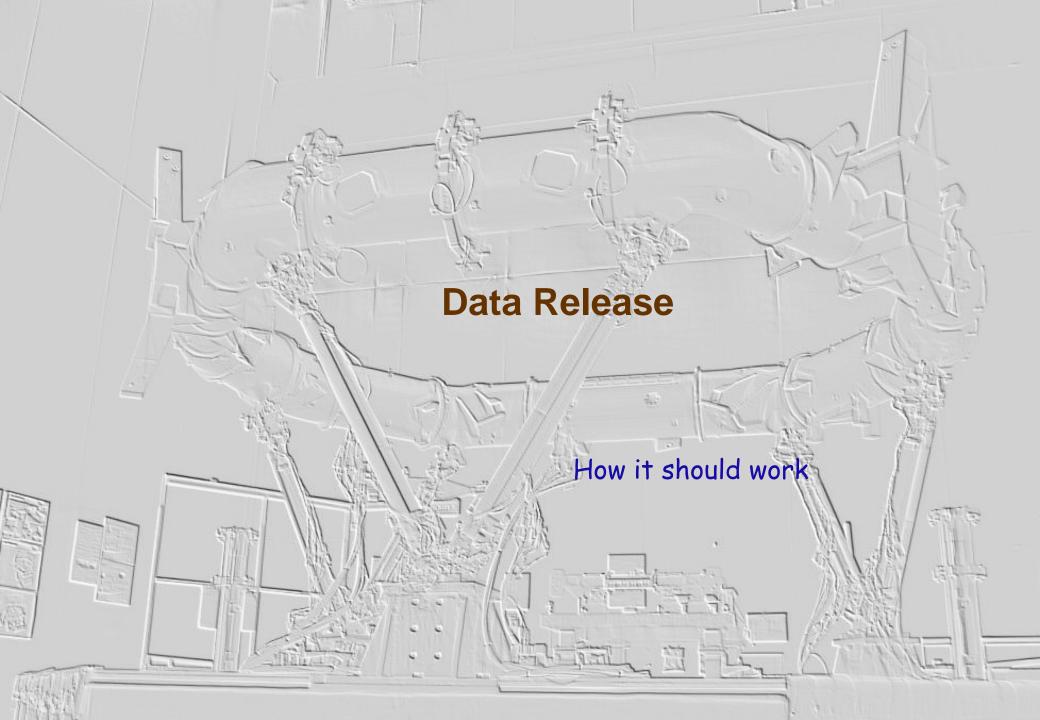
■ First dry "lift-off" successful





Credit: Arianespace

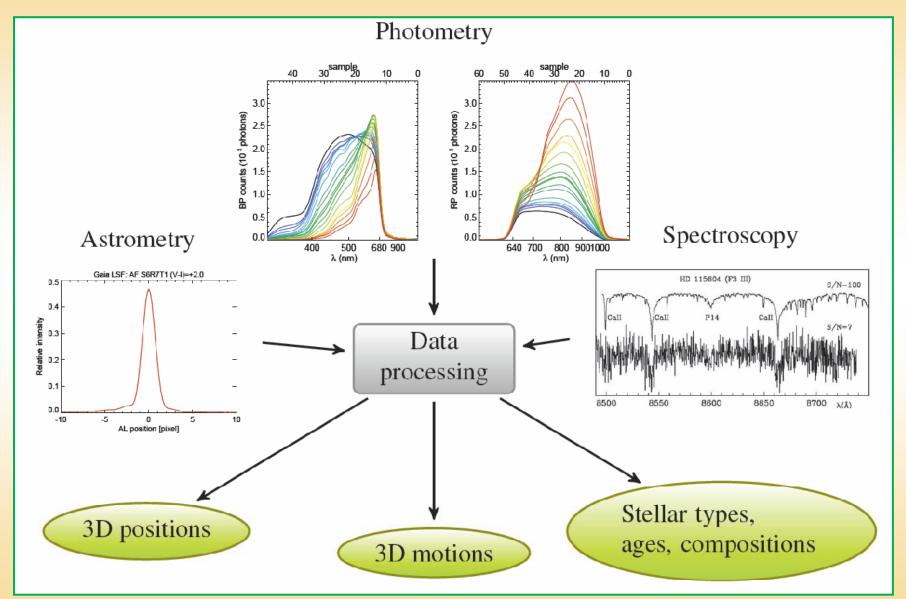
A 2011- F.





What Gaia will deliver: wide variety







Data Release



- Overall principles defined in the SMP (Science Management Plan)
 - ◆ Top level document ESA/SPC covering scientific aspects of the mission
- Intermediate results should be communicated to the Community
 - calibration will be still in imperfect state
 - feedback is expected to improve the final catalogue
- Intermediate catalogues should not delay the final catalogue
- There will be at least several intermediate coordinated release
 - this will include a first astrometric catalogue and integrated photometry
 - global accuracy will be already remarkable
 - but non statistical errors may lie everywhere
 - generic error (eg: $\sigma(G, \alpha, \delta)$) may replace source level error



Possible content of the early releases



- A position catalogue (mas precision) and G mag, when 90% of the sky is covered
 - 6-8 month of data
- The Hundred Thousand Proper Motion catalogue based on Hipparcos and Gaia positions (~ 6-8 month of data)
- Special release for the ecliptic pole region observed in the Ecliptic Pole Scanning Law
- A 5-parameter astrometric solution of astrometrically well behaved stars when it can be done for at least 90% of the sky
 - possible early results for fundamental physics parameters
- BP/RP spectrophotometric data when 5% calibration accuracy has been reached
- Mean radial velocities for stars bright enough for single epoch Rv determination with sufficient epochs and 90% of the sky
- High resolution mean spectra for mag <10 and S/N at least 50 when 90% of the sky covered
- CU defined releases like Solar System astrometry, non single stars, Epoch variability data

40



Science Alert Mode & Release

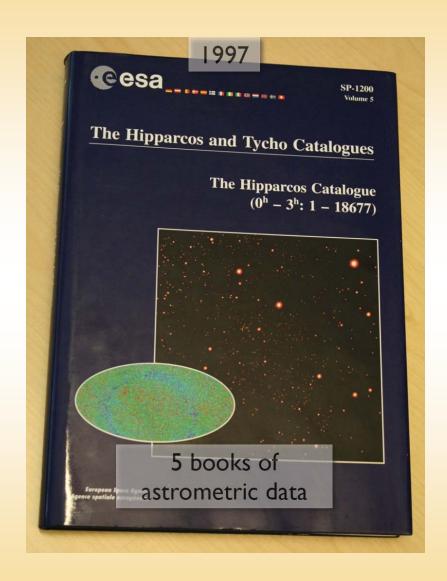


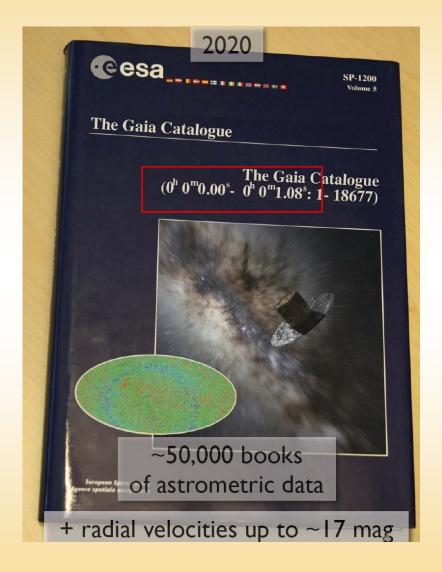
- Gaia has a built-in science alert mode:
 - Science data that would have little or no value without quick ground-based follow up
- Astrometry, Photometry and Spectroscopy could be the source of a Gaia Alert
 - a transient photometric/spectro event evidenced in the Gaia data,
 - or a fast-moving solar system object without known orbit.
 - but without possible monitoring by the Spacecraft
- Gaia releases the alert to the science community
- Immediate follow-up needs the participation of the community



Gaia Catalogue:









Catalogue Access



- Intermediate and final release will be accessed on-line
- A dedicated CU (Coordination Unit) will be set up soon
 - it will be formed sometime in 2012
 - a dedicated AO will be released by ESA
 - * An embryonic version is in place with the GAP group in the DPAC
 - GAP: Gaia Archive Preparation
 - It will develop all the necessary data mining tools to handle requests
- The actual tasks are in the process of early definition
 - requirements and specification document drafted
- Funding (not included in the current DPAC) being discussed with the national agencies
- large Expressions of Interest received from groups or countries

