GAIA NSL and orbit determination of short period binaries and planets

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Importance of the period search

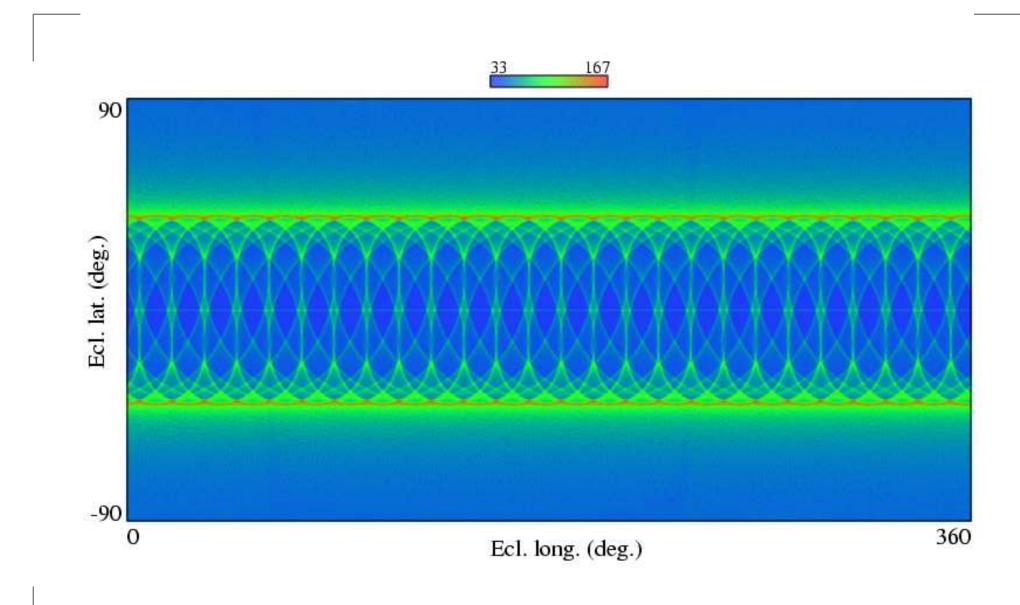
Several workgroups need a good period finder:

- DMS/PS: essential first step in astrometric orbit fitting
- DMS:essential first step in spectroscopic orbit fitting
- Variability: light curve analysis

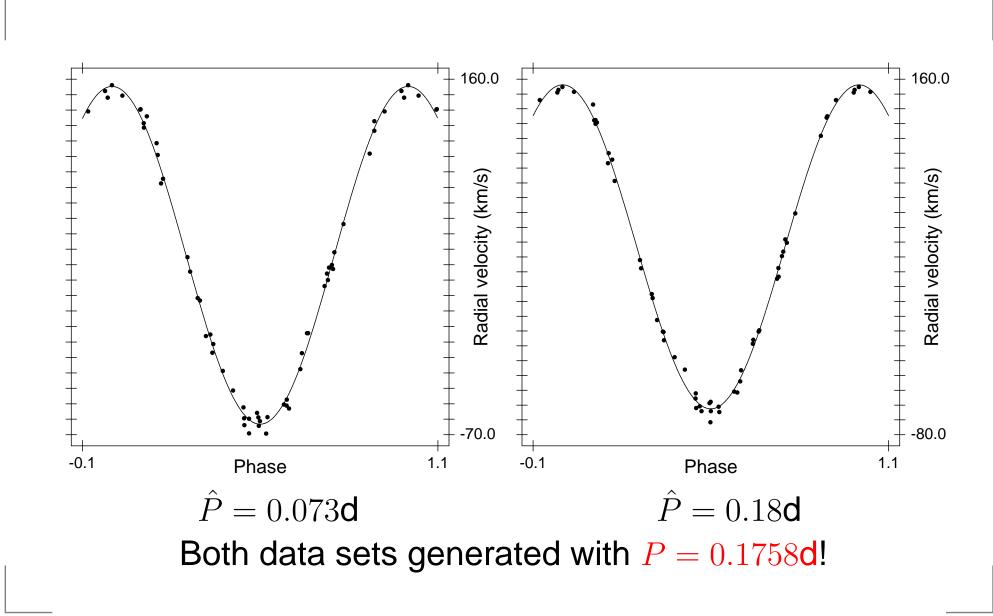
Tools:

- Scargle-like periodogram: fast but assumes a unique sine curve in the signal
- PDM-like: no assumption about the model but data must be sorted at each iteration
- Deeming (Fourier-like): no assumption about the signal, each periods could be detected in combination with the spectral window.

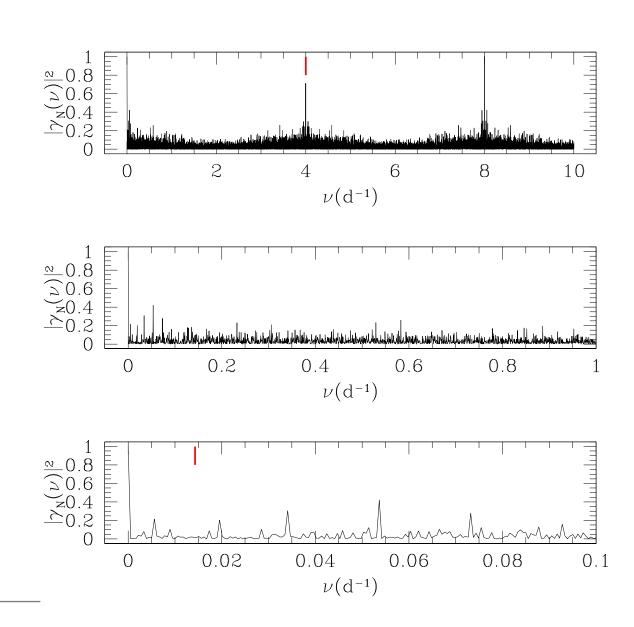
Number of observations



Which one is right?

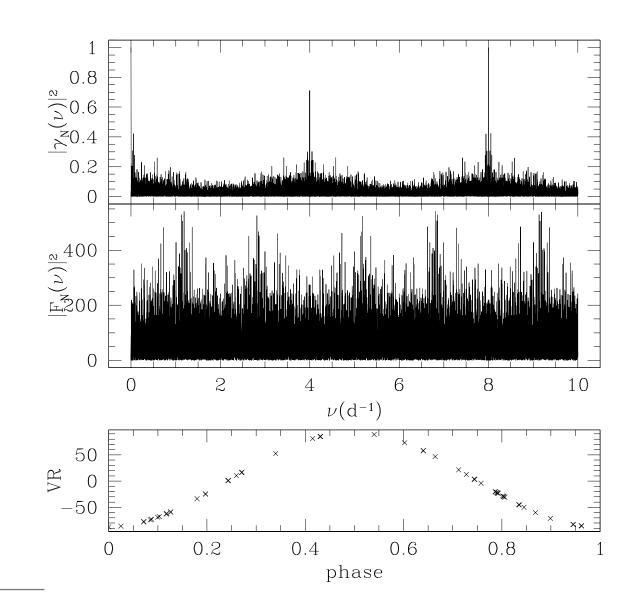


Example of a NSL spectral window



- Rotation period
 (6h) present in
 the time series
- No trace of the precession period (70d).

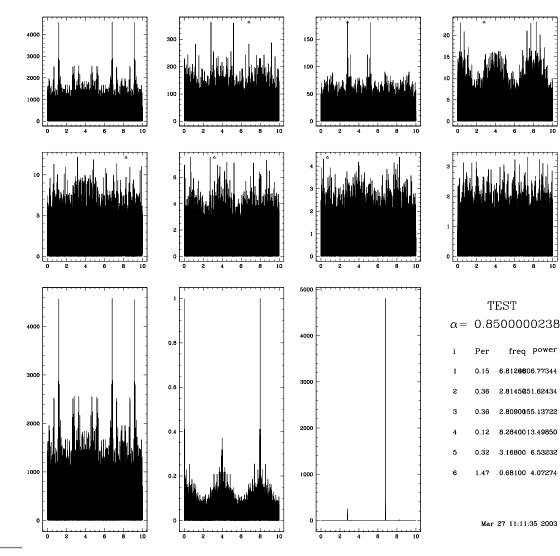
Deeming: power spectrum



High S/N data but, still, tons of peaks in the power spectrum!

Cleaning the power spectrum

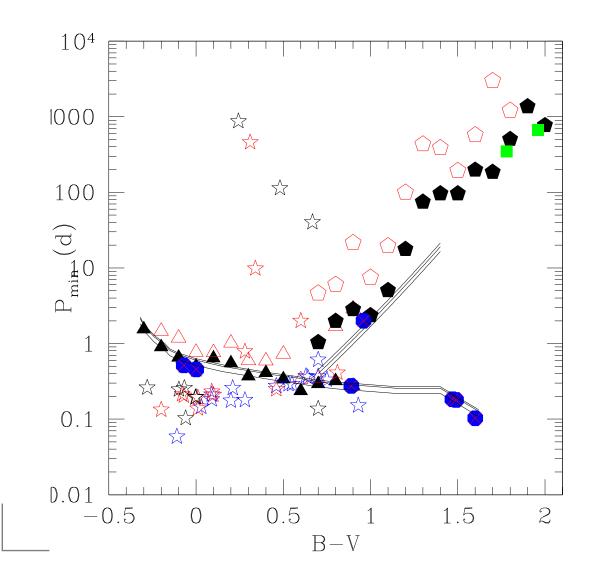
(P. Bartholdi, priv. comm.)



Idea: recovering the fundamental frequencies of the signal by removing the frequencies present in the spectral window

HR diag. and orbital period

Based on S_{B^9}



Roche lobe filling prevents short period binaries.

Position in the HR diag can limit the range of periods that need to be investigated.

Conclusion

- Even though 2 consecutive observations can be much closer than 6h, the Nyquist frequency is close to 4d⁻¹.
- One should not look for any period shorter than 6h.
- One ignores some semi-detached systems (e.g. WD+MS). According to S_{B^9} , 3.5% of the systems have orbital period shorter than 6h.
- One could break the regularity of the NSL by adding a ground-based data (e.g. from RAVE).

Which one is right?

