

# Scanning law and short period spectroscopic binaries

*How long before an orbit?*

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# Simulated data

Sample 1 (570 systems, 1710 simulations)

- 30 systems ( $w/\alpha$ ,  $\delta$ ,  $e$ ,  $\omega_A$ ,  $T$  randomly generated,  $K_A = 10$ ,  $\sigma_V = 1$ );
- $P_d \in [1, 10]$  (step:  $0.5d$ );
- Solution after 500, 1000, & 1500 days.

Sample 2 (2160 systems, 62640 simulations)

- 30 systems ( $w/\alpha$ ,  $\delta$ ,  $e$ ,  $\omega_A$ ,  $T$  randomly generated,  $K_A = 10$ ,  $\sigma_V = 1$ );
- $P_d \in [10, 730]$  (step:  $10d$ );
- Solution after 100, 150, ..., 1500 days.

# Period determination

Among the several methods for period determination, even for unevenly sampled data (e.g. Stellingwerf 1978, Scargle 1982) we give preference to Horne & Baliunas (1986):

$$P_X(\omega) = \frac{1}{2\sigma^2} \left\{ \frac{\left[ \sum_{j=1}^{N_o} (X(t_j) - \bar{X}) \cos(\omega(t_j - \tau)) \right]^2}{\sum_{j=1}^{N_o} \cos^2(\omega(t_j - \tau))} + \frac{\left[ \sum_{j=1}^{N_o} (X(t_j) - \bar{X}) \sin(\omega(t_j - \tau)) \right]^2}{\sum_{j=1}^{N_o} \sin^2(\omega(t_j - \tau))} \right\}$$

$$\text{where } \tan(2\omega\tau) = \left( \sum_{j=1}^{N_o} \sin(2\omega t_j) \right) / \left( \sum_{j=1}^{N_o} \cos(2\omega t_j) \right)$$

# Frequency range

$P_X(\omega)$  is evaluated at frequency  $\nu (= \omega/2\pi)$  in  $[\Delta T^{-1}, \nu_{Ny}]$  where  $\nu_{Ny}$  is the Nyquist frequency (Eyer & Bartholdi 1999):

$$\nu_{Ny} = \frac{1}{2p}$$

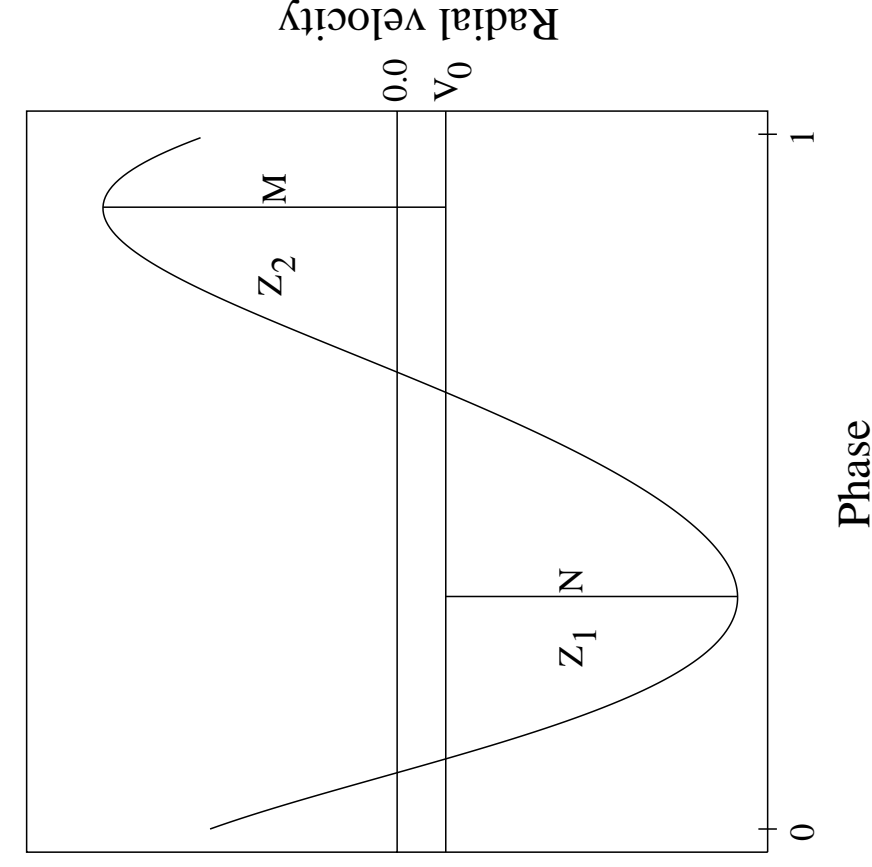
where  $p$  is the largest value such that  $\forall t_j, t_j = t_1 + n_j p$  with  $n_j \in N$ .

Step:  $\frac{1}{\Delta T}$

Probability of false detection  $F = 1 - [1 - e^{-z}]^{N_i}$  where

$$N_i \approx -6.4 + 1.2N_o + 0.00098N_o^2$$

# Initial solution: Lehmann-Filhés



The curve is replaced by the polygonal contour of the data.

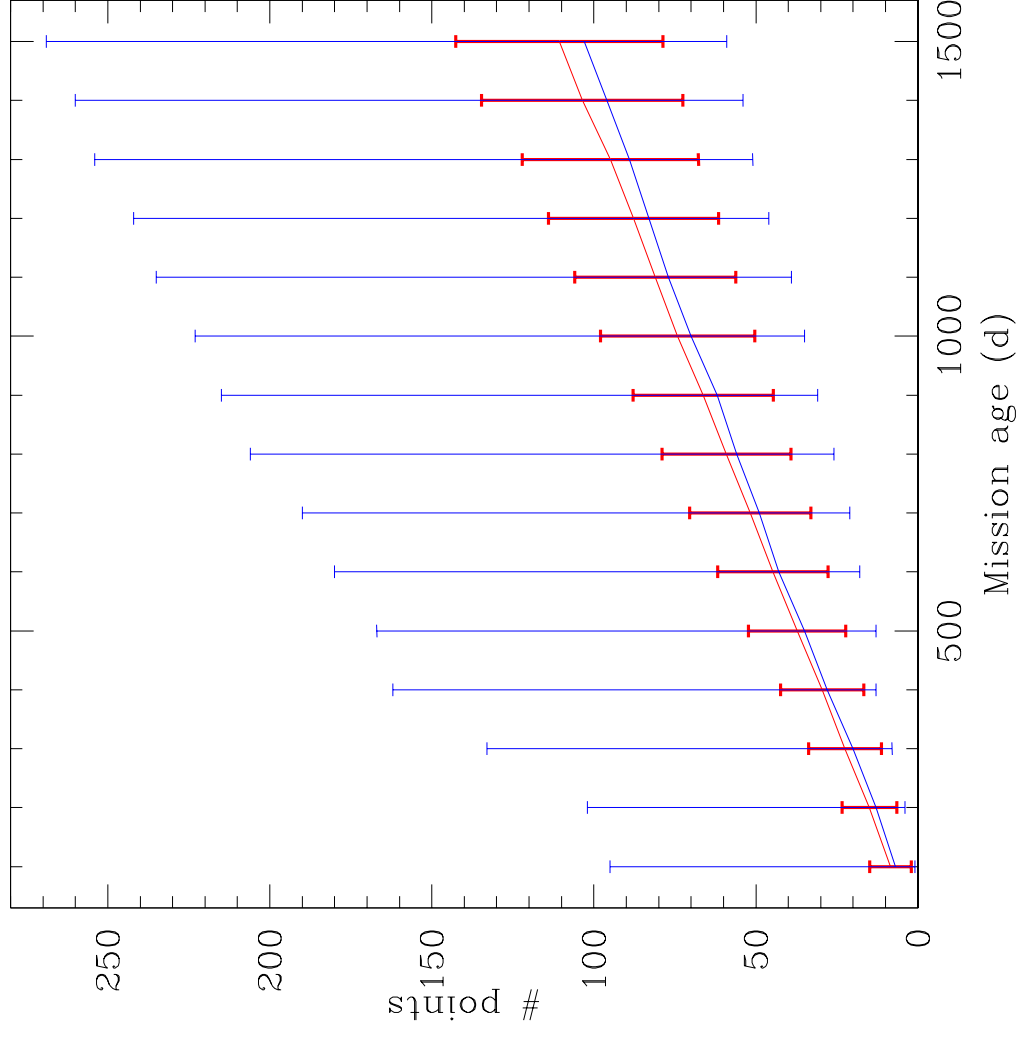
$V_0$  is such that the areas (trapezoidal rule) above and below  $V = V_0$  are equal.

$$e \cos(\omega) = \frac{M - N}{M + N}$$

$$e \sin(\omega) = \frac{2\sqrt{MN} Z_2 - Z_1}{M + N Z_2 + Z_1}$$

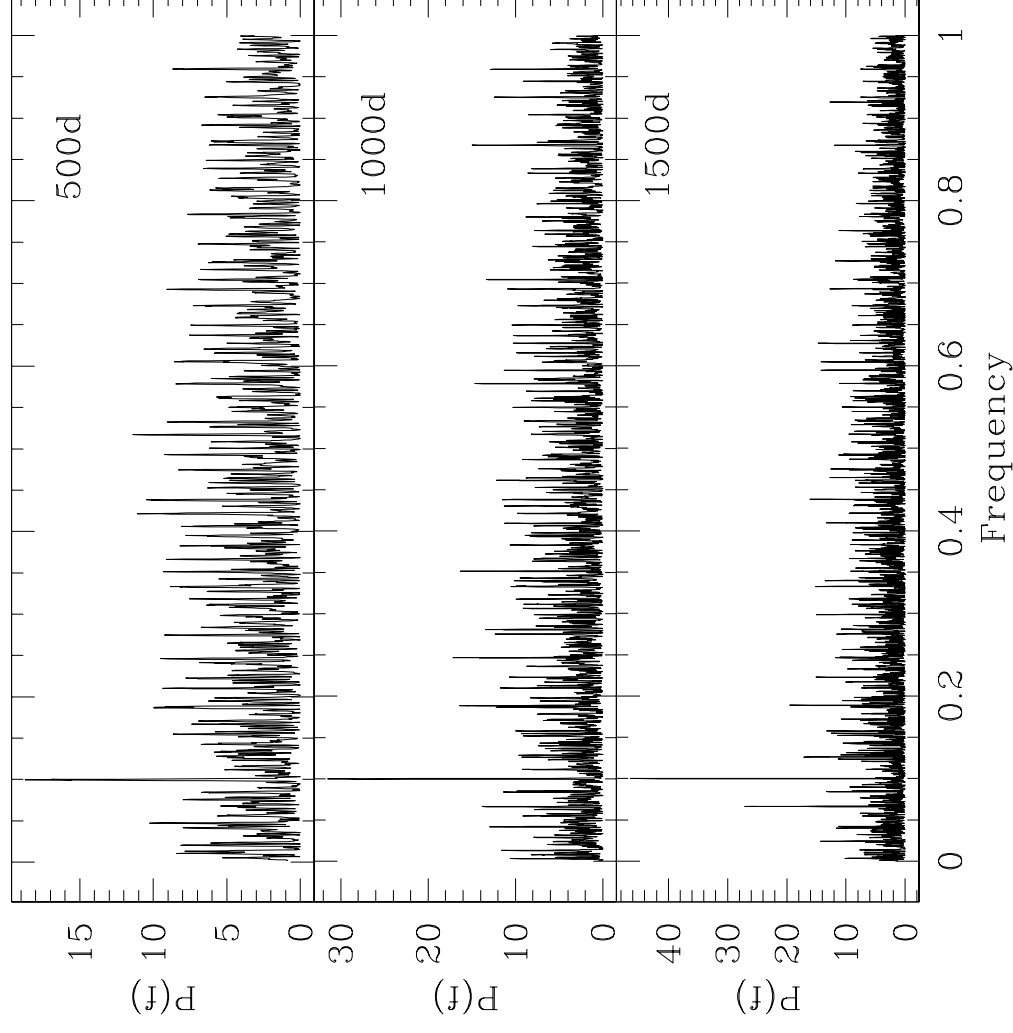
# Sample size

From sample 2 (2160 systems)



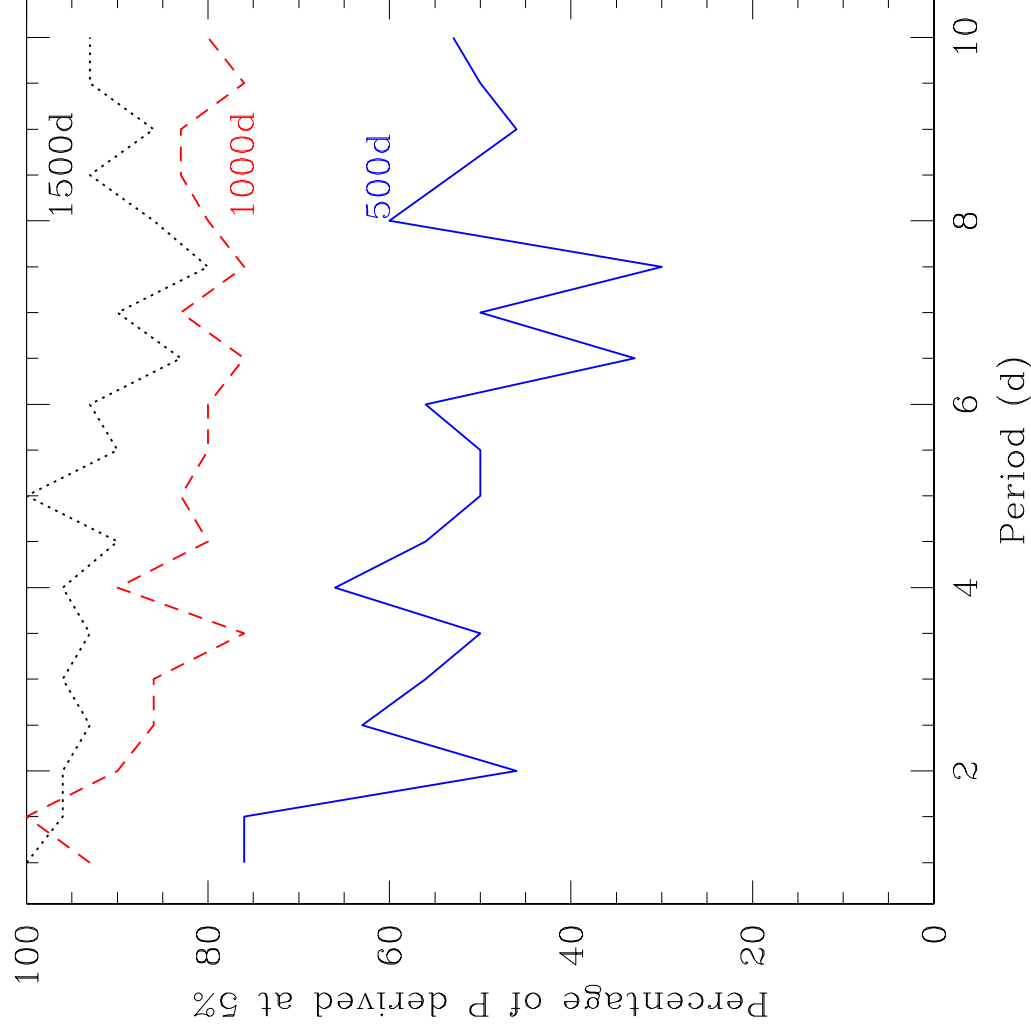
# Power spectrum at 3 epochs

Sample 2



# Derived period

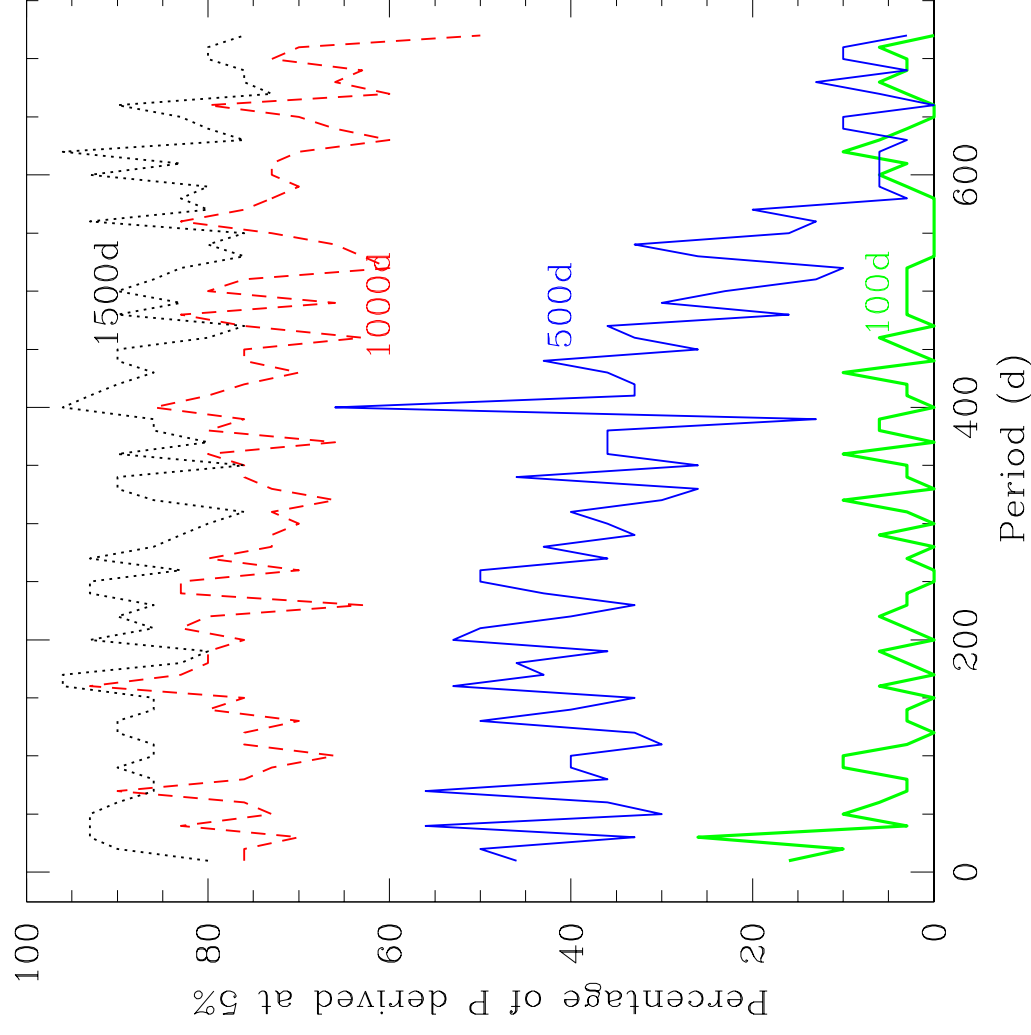
Sample 1





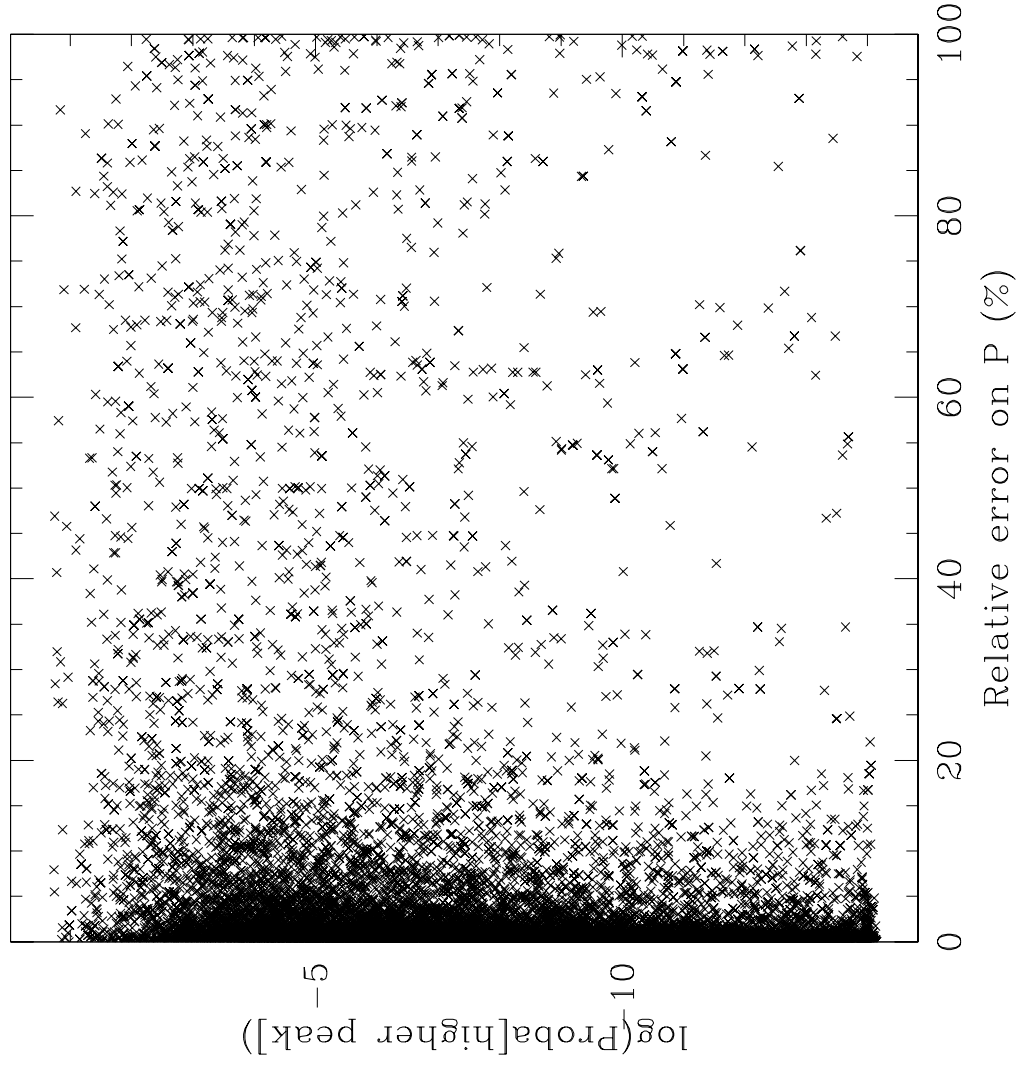
# Derived period (cnt.)

Sample 2



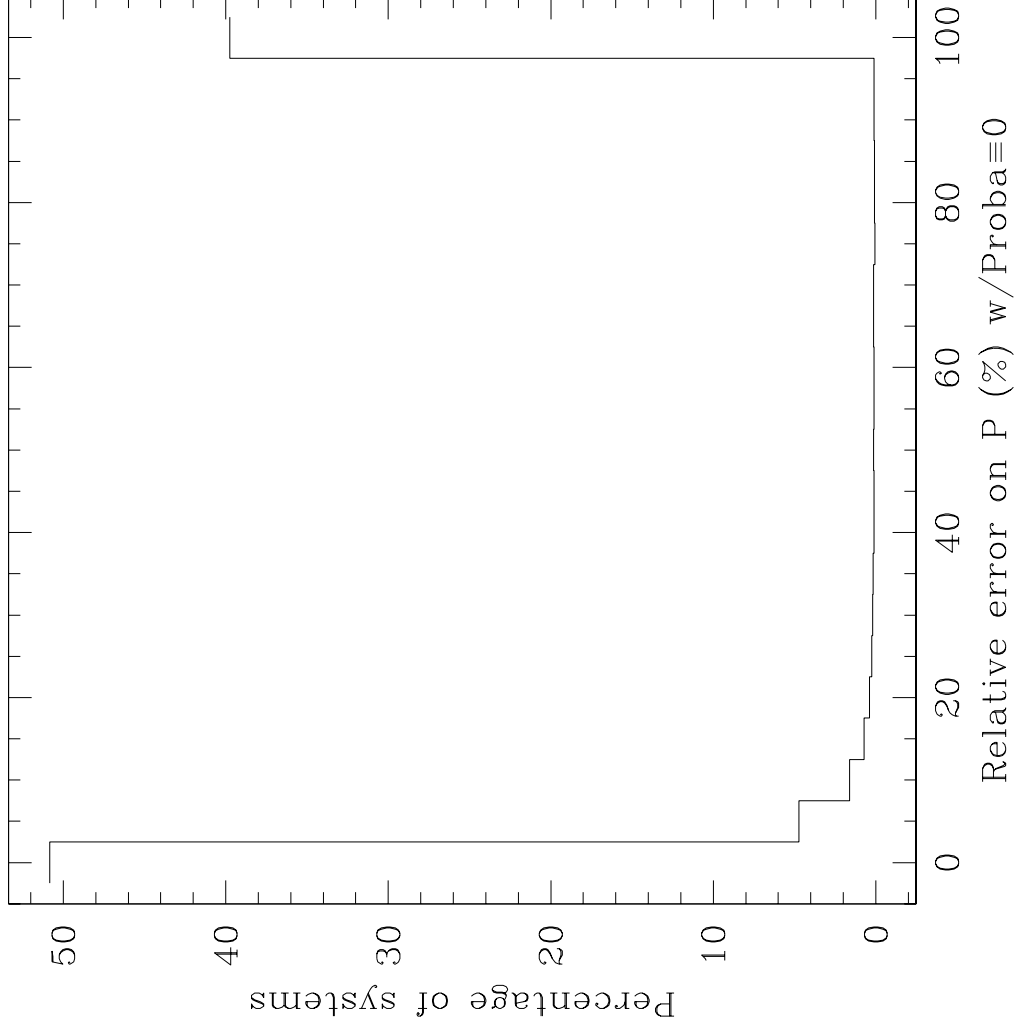
# Period vs. F

Sample 2



# Period when $F=0$

Sample 2



# Conclusions

- Periodogram: quick and easy way to get a period. Faster implementation. Combination with other methods?
- Although  $S/N = 10$ , the relevance of highest peak looks questionable,  $F$  does not help!
- The scanning law often prevents the short periods (a few days) from being identified rapidly.
- The percentage of right **orbits** is even **lower**.

The coordinates should be left out of the simulation  
↗ percentage vs. location on the sky.