First Look Overview - Relevance for CU6

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1st Meeting of the Gaia CU6 "Spectroscopic Processing"



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Introduction	First Look	CU3 Data Flow	Science Quick Look	Detailed First Look	Remarks and Conclusion

Outline



- 2 First Look
 - Overall QL/FL Scheme
 - "Lunch" Scheme
 - ODAS: ODIS/Ring Solution
- 3 CU3 Data Flow
 - Science Quick Look
- Detailed First Look
- 6 Remarks and Conclusion



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Introduction First Look CU3 Data Flow Science Quick Look Detailed First Look Remarks and Conclusion ● 0000

The Basic Problem

Proper functioning of all elements

Problem

It is difficult to immediately assess proper functioning of all elements at the μ as level!

Specific Problem

 A full solution of the astrometric problems is possible only after more than six months!



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Introduction First Look CU3 Data Flow Science Quick Look Detailed First Look Remarks and Conclusion • 0000

The Basic Problem

Proper functioning of all elements

Problem

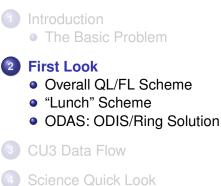
It is difficult to immediately assess proper functioning of all elements at the μ as level!

Importance

 However, this is very important, else risk of losing many months of data and mission time!



Outline



- 5 Detailed First Look
- 6 Remarks and Conclusion



Science Quick Look

Detailed First Look

Remarks and Conclusion

Overall QL/FL Scheme

Different Levels of QL/FL

- Quick-Look (QL)
- Science Quick Look (ScQL)
 - ScQL monitor and evaluator
- Initial data treatment (IDT)
 - Ingestion
 - Cross-matching
 - Image parameter estimations
- First Look Preprocessing (FLP)
 - ODAS (One-Day Astrometric Solution)
 - PSF/LSF
 - Photometry
 - RVS, CCD
- Oetailed First Look (DFL)
 - First-Look monitor
 - First-Look evaluator

Responsible

ESOC

realtime

HK data



Science Quick Look

Detailed First Look

Remarks and Conclusion

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- Oetailed First Look (DFL)
 - First-Look monitor
 - First-Look evaluator

Responsible

ESOC

- + FL Task support
 - < 30 min or realtime
 - $\bullet \approx 10 \, \text{mas}$

S/C HK
+Science HK

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Science Quick Look

Detailed First Look

Remarks and Conclusion

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- RVS, CCD
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 - First-Look evaluator

Responsible

Core Processing

few hours



Science Quick Look

Detailed First Look

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First Look Preprocessing (FLP)

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- Photometry
- RVS, CCD
- Detailed First Look (DFL)
 - First-Look monitor
 - First-Look evaluator

Responsible

FL Task

- daily
- \approx 10 μ as

HK

+science data



Science Quick Look

Detailed First Look

Remarks and Conclusion

Overall QL/FL Scheme

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- Photometry
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Responsible

FL Task

- daily
- pprox 10 μ as

HK

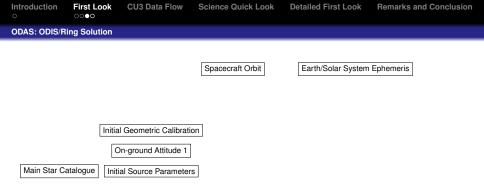
+science data

Introduction O		CU3 Data Flow	Science Quick Look	Detailed First Look	Remarks and Conclusion
"Lunch" Scher	ne				

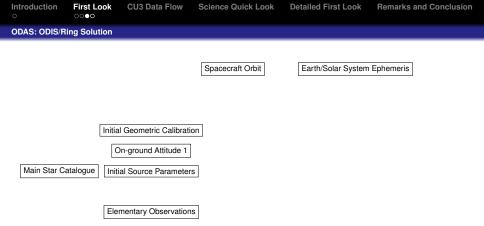
Responsibilities within the RVS First Look tasks

Task	Definition	Development	Integration	Operation
QL	ESOC (input from industry, CU3, CU6)	ESOC	ESOC	ESOC
ScQL	CU3, CU6, ESOC, industry	under discussion	ESOC	ESOC
DFL Monitor (using IDT)	CU6, CU3	CU6	ESAC	ESAC (automatic)
DFL Evaluator (using IDT)	CU6, CU3	CU6	ZAH (CNES)	ZAH (mirrored \rightarrow CU6)
DFL Monitor	CU6, CU3	CU6	CNES	CNES (automatic)
DFL Evaluator	CU6, CU3	CU6	ZAH (CNES)	ZAH (mirrored \rightarrow CU6)

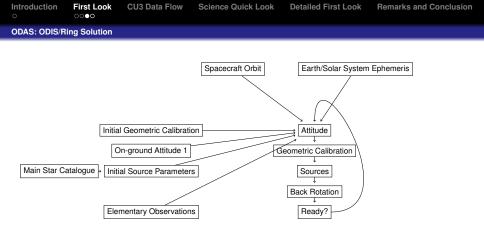




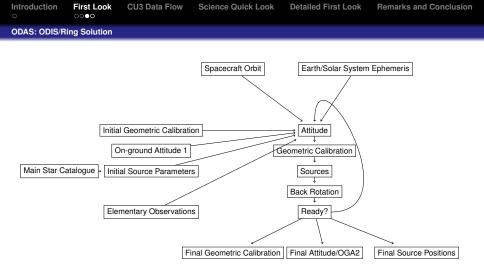




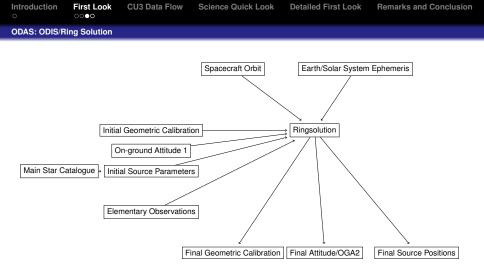




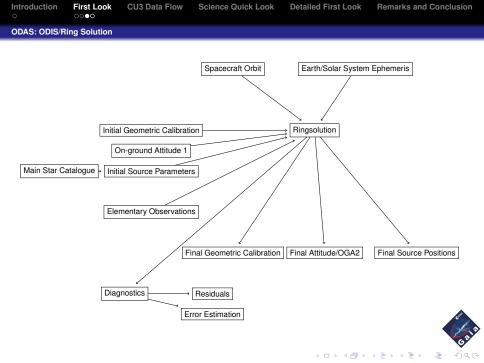


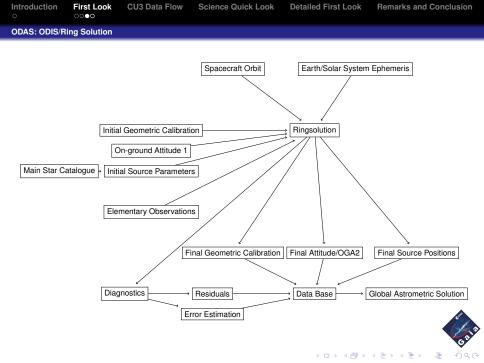












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ODAS: ODIS/R	ing Solution				

ODAS Results

- Starting from large errors (150000 μas) in geometric calibration, source positions, and attitude we obtain final errors in
 - calibration: $< 1\mu$ as (along scan), $< 10\mu$ as (across scan)
 - source positions: $\approx 80\mu as$ (along scan), $\approx 900\mu as$ (across scan)
 - attitude: \approx 20 μ as (along scan), \approx 600 μ as (across scan)
- Convergence is very slow in ODIS, but can be accelerated
- Ring Solution: Direct solution in one/two steps, same result as with ODIS but allows a direct evaluation of the statistical errors
- The solution is extremely valuable as input for the Global Astrometric Solution or any alternative of it (e.g. Ring2Sphere)



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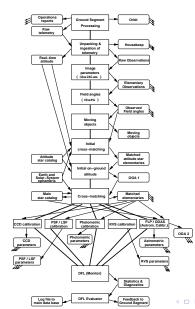
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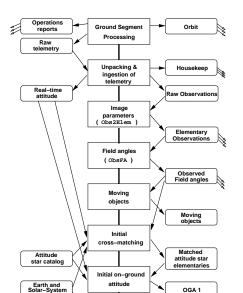


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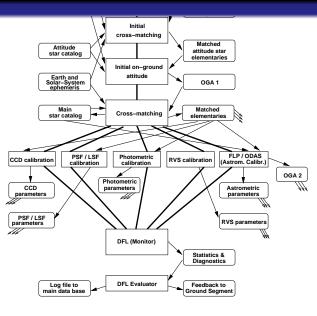




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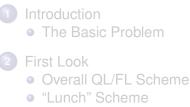






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General idea

The Science Quick Look looks at a tiny part of the science data (Science HK data)

Selection

• For the SM/AF/BP/RP chain: a randomly selected tiny sample of windows with stars 13 < *G* < 16



General idea

The Science Quick Look looks at a tiny part of the science data (Science HK data)

Amount

About 1/2000 of all data



General idea

The Science Quick Look looks at a tiny part of the science data (Science HK data)

Typical Diagnostics

Are the stars centered in the window?



General idea

The Science Quick Look looks at a tiny part of the science data (Science HK data)

Typical Diagnostics

 How does the width of an image compare to the assumed PSF/LSF (color dependent)



General idea

The Science Quick Look looks at a tiny part of the science data (Science HK data)

Typical Diagnostics

- What is the background, how large is the noise?
- Does the onboard background match the value computed on ground?



General idea

The Science Quick Look looks at a tiny part of the science data (Science HK data)

RV ScQL?

 What kind of diagnostics are possible with the RV windows for single transits (without knowledge of the star's identity!) or a statistic thereof?



General idea

The Science Quick Look looks at a tiny part of the science data (Science HK data)

RV ScQL?

• What kind of diagnostics are possible with flags/counters integrated into the on-board software?



General idea

The Science Quick Look looks at a tiny part of the science data (Science HK data)

RV ScQL?

 Onboard-diagnostics: number of removed cosmics (one example from Katz et al., GAIA-C6-TN-OPM-DK-001-2)



General idea

The Science Quick Look looks at a tiny part of the science data (Science HK data)

RV ScQL?

 Window-diagnostics: Is the number of photons in the spectra consitent with the brightness/color type?



General idea

The Science Quick Look looks at a tiny part of the science data (Science HK data)

RV ScQL?

- Do you need the information from AF/BP/RP instruments?
- In this case one must define that the same objects are selected for ScQL?



General idea

The Science Quick Look looks at a tiny part of the science data (Science HK data)

Onboard diagnostics

- Please provide us with input to coordinate this!
- Coordination between GWP-S-610-00000 (Desert with Mignot, Hébrard, Lecavelier) and GWP-M-355-00000 (Jordan/Biermann et al.)



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Science Quick Look

General idea

The Science Quick Look looks at a tiny part of the science data (Science HK data)

Science Quick Look

• Please keep everything as simple as possible!



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Remarks and Conclusion



General idea

First Look Preprocessing (calibration, etc...) + other diagnostic on a time scale of up to one day

Astrometric diagnostics

Breakdown into time intervals.



General idea

First Look Preprocessing (calibration, etc...) + other diagnostic on a time scale of up to one day

Astrometric diagnostics

Is the attitude stable?



General idea

First Look Preprocessing (calibration, etc...) + other diagnostic on a time scale of up to one day

Astrometric diagnostics

Is the geometric calibration stable?



General idea

First Look Preprocessing (calibration, etc...) + other diagnostic on a time scale of up to one day

Astrometric diagnostics

• Are the residuals from ODAS compatible with the desired accuracy of the elementary measurements?



General idea

First Look Preprocessing (calibration, etc...) + other diagnostic on a time scale of up to one day

Non-astrometric Diagnostics

 These have to be provided by other tasks (photometry, RVS, CCDs, ...)!!



Non-calibrational diagnostics

Diagnostics not directly connected to calibrational issues (like ODAs in the case of astrometry)

Non-calibrational diagnostics

• The DFL uses fainter stars then QL or ScQL.



Non-calibrational diagnostics

Diagnostics not directly connected to calibrational issues (like ODAs in the case of astrometry)

Non-calibrational diagnostics

 Is the number of windows as large as expected for different magnitudes, types of stars ...?



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Non-calibrational diagnostics

Diagnostics not directly connected to calibrational issues (like ODAs in the case of astrometry)

Non-calibrational diagnostics

• Are the centroids of the images in the center of the windows?



Non-calibrational diagnostics

Diagnostics not directly connected to calibrational issues (like ODAs in the case of astrometry)

Non-calibrational diagnostics

• Are the images focused (PSF, LSF)?



Non-calibrational diagnostics

Diagnostics not directly connected to calibrational issues (like ODAs in the case of astrometry)

Non-calibrational diagnostics

Are there CCD defects?



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First Look

The First Look checks the overall functioning of Gaia on many different levels of accuracy and by different means

Breakdown

QL, ScQL, DFL



First Look

The First Look checks the overall functioning of Gaia on many different levels of accuracy and by different means

Importance

 Its provides important diagnostics for astrometry, photometry, RV, etc...



First Look

The First Look checks the overall functioning of Gaia on many different levels of accuracy and by different means

Importance

 Moreover, it is an essential ingredient of the overall data flow



First Look

The First Look checks the overall functioning of Gaia on many different levels of accuracy and by different means

Importance

 (At least) on the astrometric side it provides important initial values for the following Global Data Analysis (GIS, Ring2Sphere, ...)



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First Look

The First Look checks the overall functioning of Gaia on many different levels of accuracy and by different means

Important!

• The FL team need input from other tasks (IDT, Photometry, RV, ...) to perform its assignment!



First Look

The First Look checks the overall functioning of Gaia on many different levels of accuracy and by different means

Important!

• The FL team is the single contact point between science and the ground segment!



First Look

The First Look checks the overall functioning of Gaia on many different levels of accuracy and by different means

Important!

ullet \to there should be no separate FL for the different tasks

